

Many people involved...

Lukasz Wyrzykowski Guy Rixon Heather Campbell Thomas Wevers Sergey Koposov

Simon Hodgkin Nadia Blagorodnova Peter Jonker Sjoert Van Velzen Nic Walton Gerry Gilmore



See

http://www.ast.cam.ac.uk/ioa/wikis/gsawgwiki/index.php/Working_groups

gaia.ac.uk/selected-gaia-science-alerts



Alerts

The table can by sorted by Name, UTC timestamp, RA, Dec and AlertMag - click column heading to

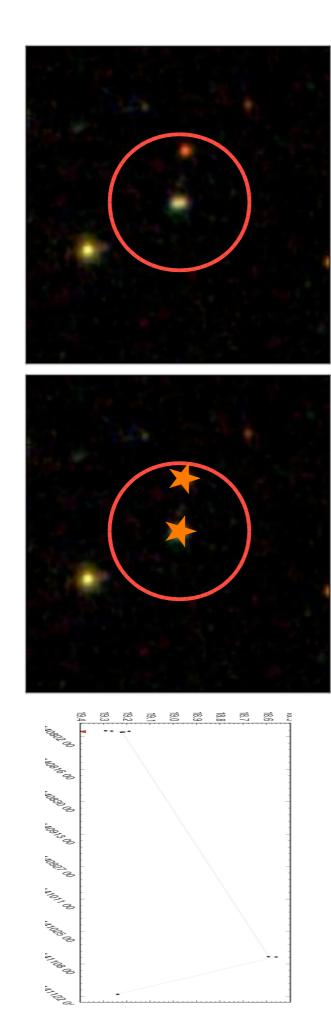
Gaia I 4adh	Gaia I 4adi	Gaia I 4adj	Name -
Gaia14adh 2014-11-05	2014-11-07 01:05:09	2014-11-17 20:25:19	UTC timestamp
168 69436 <u>-</u> 13 88694 15 84	168.47841	356.72155 23.14965 18.91	RA
-13 88694	168.47841 -23.01221 18.77	23.14965	Dec
15 84	18.77		AlertMag
18 45	19.62	19.53	HistMag
0 23	0.07	0.08	AlertMag HistMag HistStdDev Class
IInknown	unknown	unknown	Class
Blue in BP/RP and SDSS	Fading source on top of 2MASS Galaxy (offset from bulge)	Early type galaxy with an extra source. Could be a foreground star. Imaging target	Comment
25 Nov 2014	2 Dec 2014, 13:55	2 Dec 2014, 15:18	Published

Methodology

- Search around known SDSS/DSS/2MASS sources

For each epoch, sum flux of all Gaia sources within 4" into "superpixel"

- Search for Δmag > 0.5 (and at >5σ) over historic magnitude
- Plus present in next visit

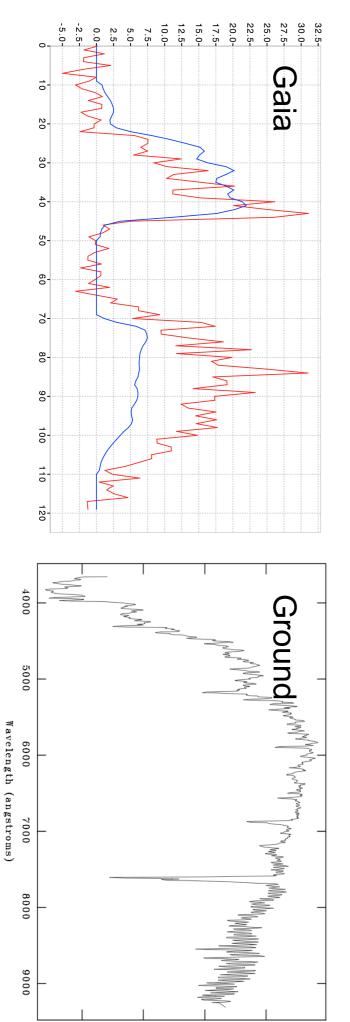


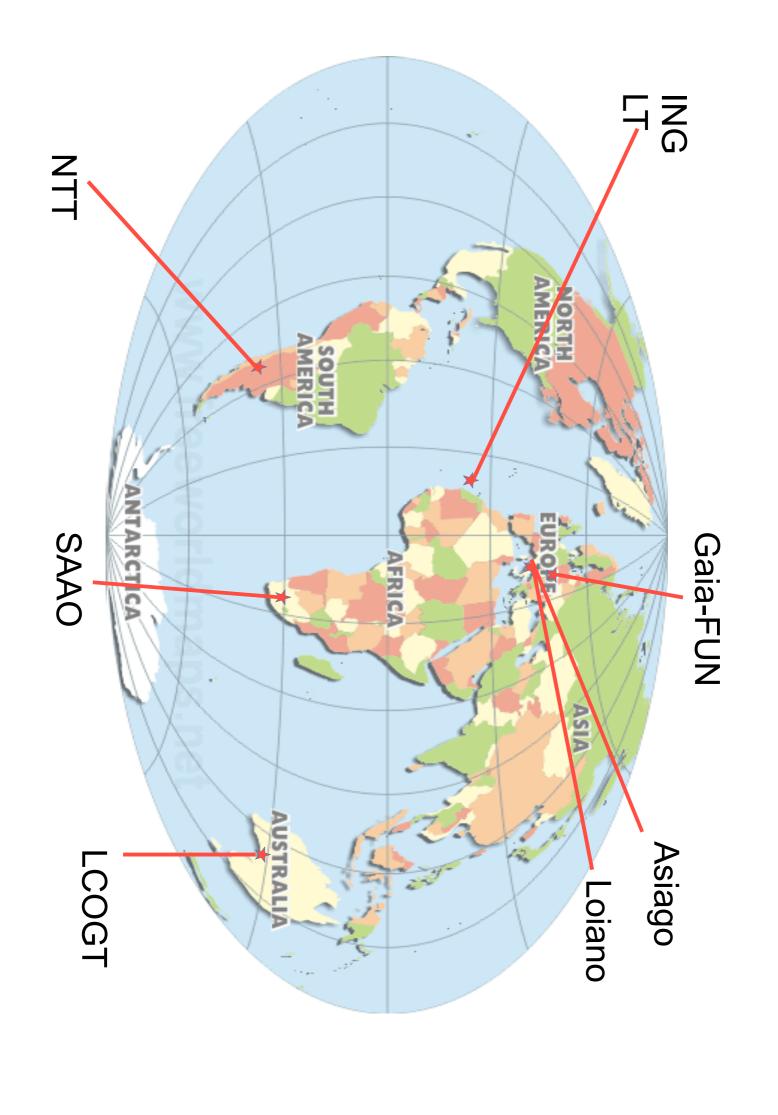
Reasons for followup

- Many transient science cases need higher cadence
- Or higher resolution spectroscopy...

Fainter targets will not get BP/RP spectra

- Validation of BP/RP spectra, training set
- Validation of alerts





Welcome to the Cambridge Photometry Calibration Server (CPCS)

Not logged in

Login into the system List of alerts (observed only)

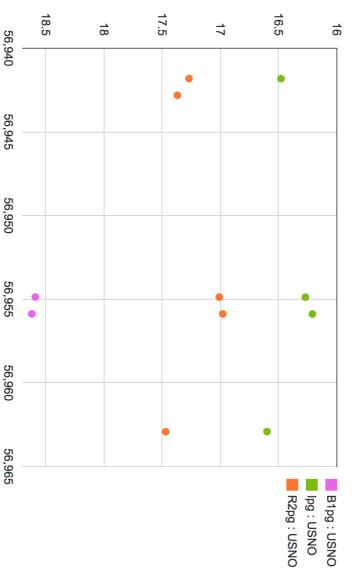
List of observatories List of followup data

Light curve of ivo://GaiaVerif14ahk

Upload new followup data

Enter new event

Delete a followup point from the system



mag

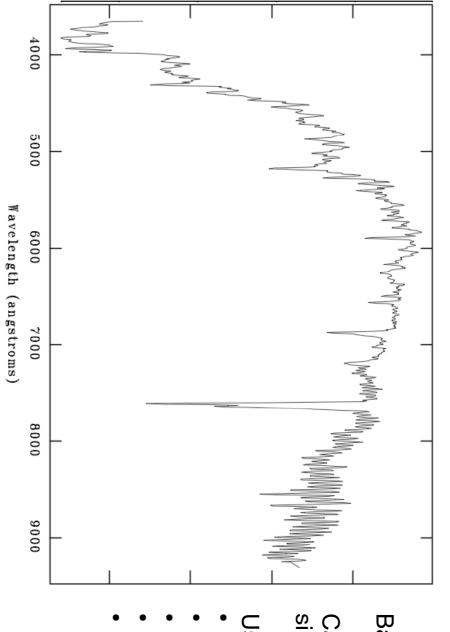
La Palma telescope time

2014B semester: Large amount of spectroscopic time during

ODOSpec
WHT+ISIS,ACAM
WHT+ISIS,ACAM
Telescope

Pipelines

\$ python int_fastspec.py r1035253.fit



Based on IRAF/Pyraf/Python

single, raw spectral frame Can be run at the telescope on a

Uses archival calibrations only

- Bias removed with overscan
- No flat-fielding Archival dispersion solution
- Wavelength shift using sky lines
- Fixed sensitivity curve

Three fast pipelines (for INT+IDS, WHT+ACAM, WHT+ISIS)

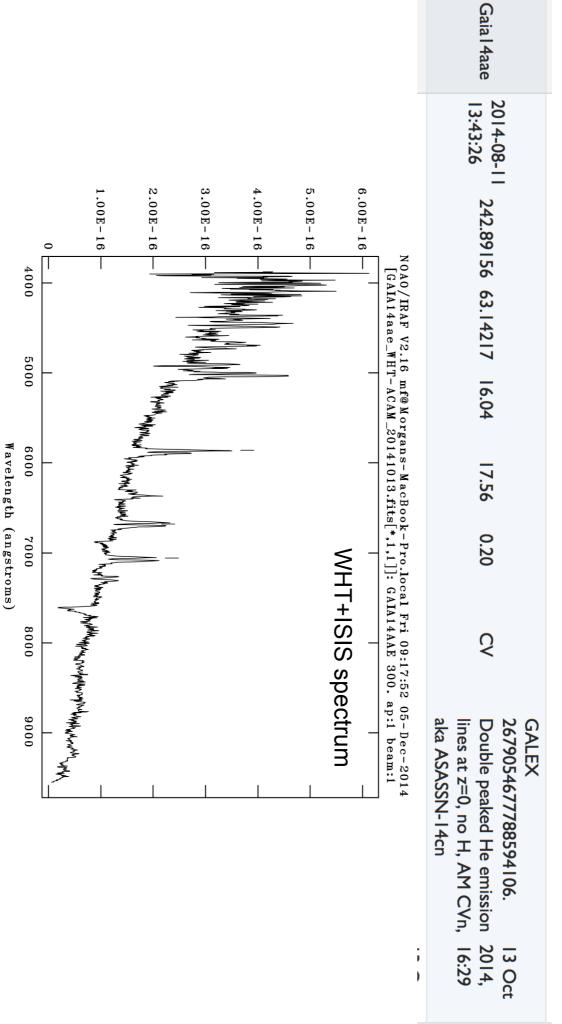
Pipelines

We are also working on a "full" pipeline (including telluric, cosmic ray corrections etc.)

and documentation) when ready. All pipelines will be released (including sources

Gaia14aae

AM CVn system seen by Gaia (independently discovered by ASASSN)



Gaia14aae



675

University), S. Komossa (MPIfR), Z. Kostrzewa-Rutkowska, S. Littlefair (University of Sheffield) A. Mahabal (Caltech), P. O'Brien (University of Leicester), P. Ochner (Padova Observatory), (University of Sheffield), K. Ilkiewicz (Warsaw University Observatory), M. Irwin (University of Davis (Liverpool John Moores), F. De Angeli (University of Cambridge), M. Dennefeld (Institut ATel #6593; G. Rixon, M. Fraser, S. Koposov, N. Blagorodnova, H. Campbell, S. T. Hodgkin, G. Observatory), A. Scholz (University of St. Andrews), M. Sitek (Warsaw University Observatory), University of Warwick), G. Busso (University of Cambridge), R. Busuttil (Open University), C. van Leeuwen, N. Walton (University of Cambridge), L. Wyrzykowski (Warsaw Observatory), D. Steeghs (University of Warwick), L. Tomasella (Padova Observatory), M. Turatto (Padova d'Astrophysique de Paris), N. Elias-Rosa (Padova Observatory), D. W. Evans (University o Wevers, P. Jonker, S. van Velzen (RU Nijmegen), S. Benetti (Padova Observatory), E. Breedt (University of Wroclaw), R. Pretorius (University of Oxford), K. Rybicki (Warsaw University (Warsaw University Observatory), B. Handzlik (Warsaw University Observatory), L. Hardy Gilmore (University of Cambridge), A. Gomboc (University of Ljubljana), A. Hamanowicz Observatory), K. Ulaczyk (Warsaw University Observatory), F. van Leeuwen (University of Riello (University of Cambridge), M. Pawlak (Warsaw University Observatory), A. Pigulski Cambridge), J. F. Jarvis (Open University), D. A. Kann (TLS Tautenburg), U. Kolb (Open Cambridge), D. Froebrich (University of Kent), B. Gaensicke (University of Warwick), G. on 19 Oct 2014; 14:48 UT 662 663 663 664 664 671 671 675 659 664

Credential Certification: Morgan Fraser (mf@ast.cam.ac.uk)

659

658

Subjects: Optical, Binary, Cataclysmic Variable, Star, Transient, Variables

Referred to by ATel #: 6626, 6640, 6641, 6646, 6750

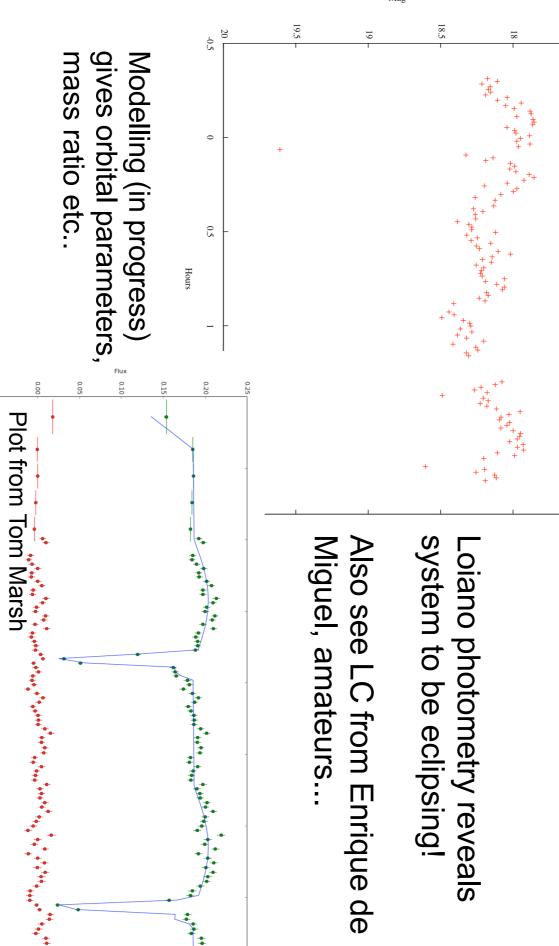
V Tweet

We report spectroscopic confirmation and classifications for reported Gaia Photometric Science Alerts (http://gaia.ac.uk/selected-gaia-science-alerts). All spectra were taken at the 4.2m William Herschel Telescope on La Palma, over the nights of 13-15 October 2014. Low to intermediate resolution spectra were taken using ISIS+R300B/R158R and ACAM+V400, and reduced and calibrated spectra are available for download (http://gaia.ac.uk/sites/default/files/file_attach/FUGA_WHT_20141013-15.tar.gz)

Name	RA	Dec	Instrument	Date of spectrum
GAIA14AAE	242.891.6	63.14217	ACAM	2014-10-13.9
GALALTANT	244.25381	62.00685	ACAM	2014-10-13.9/15.8
GAIA14AAI	23.88339	-20.40796	ACAM	2014-10-14.0
GAIA14AAT	52.76852	17.42784	ACAM	2014-10-14.0
GAIA14AAU	237.94493	16.62637	ACAM	2014-10-13.8
GAIA14ABG	262.69947	50.00445	ACAM	2014-10-13.9
GAIA14ABQ	58.59016	-11.31998	ISIS	2014-10-15.2/16.0
GAIA14ABR	59.38013	-9.65999	ACAM	2014-10-15.2
GAIA14ABY	55.51929	-22.17277	ISIS	2014-10-15.1

We acknowledge ESA Gaia (http://cosmos.esa.int/gaia), and the DPAC Photometric Science Alerts Team (http://gaia.ac.niented.gaia.ccience.alerte)

Gaia14aae



Heather Campbell (IoA) leading followup campaign

Summary

- Gaia Alerts are flowing
- Some ad-hoc solutions for now, but will get better, and on fairly short timescales
- Already finding interesting transients
- The first science from Gaia!

mf@ast.cam.ac.uk